

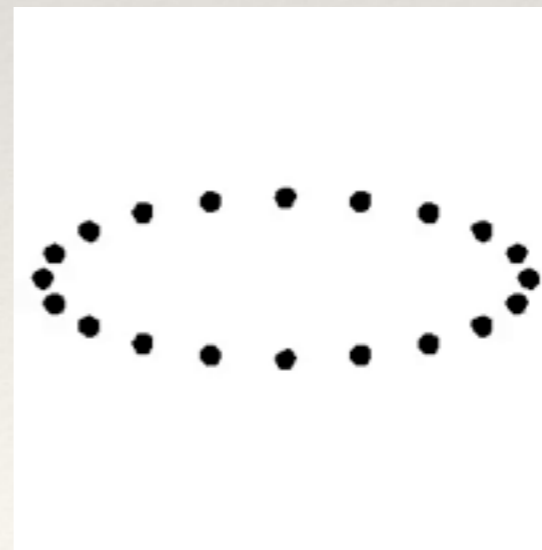
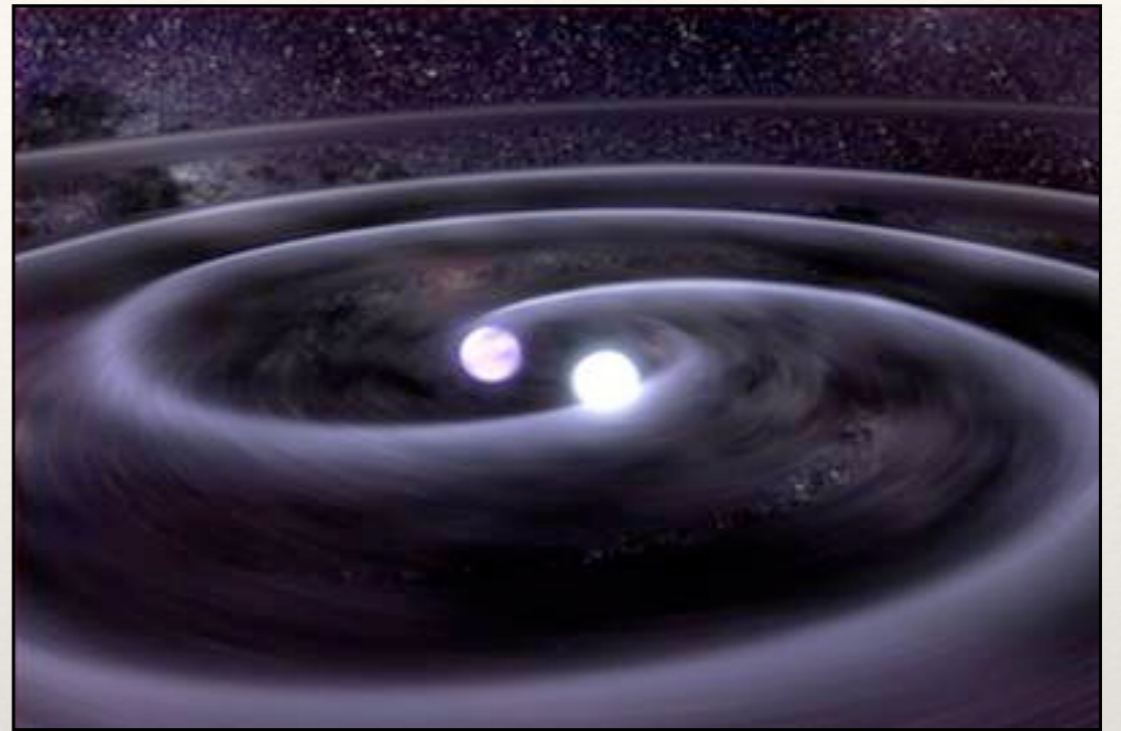


Future directions in gravitational physics

Stephen Fairhurst
Cardiff University

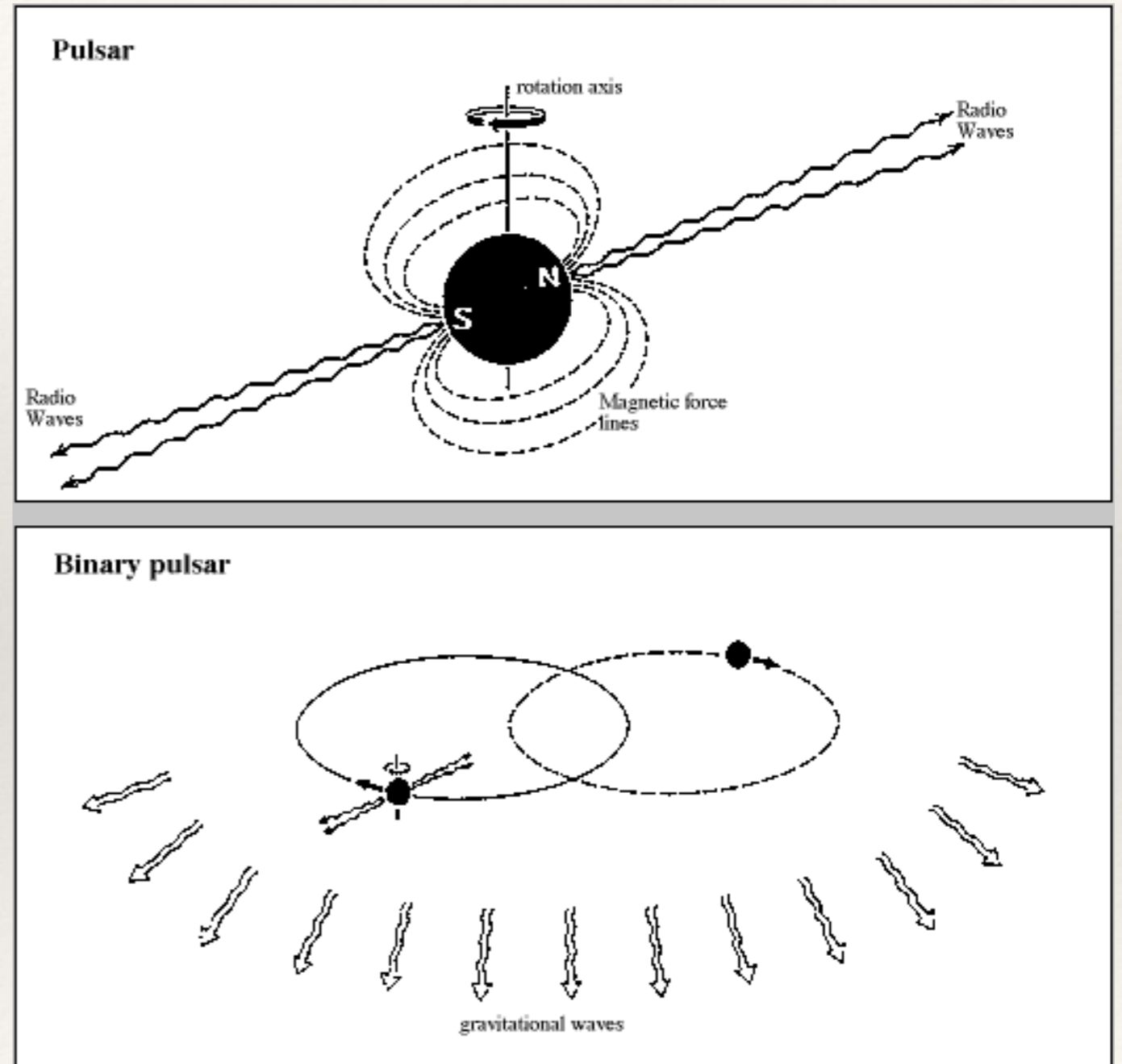
Gravitational Waves

- Gravitational waves are a feature of Einstein's general theory of relativity
- Accelerating mass emits gravitational waves
- “Ripples in space-time” propagate away from source at the speed of light



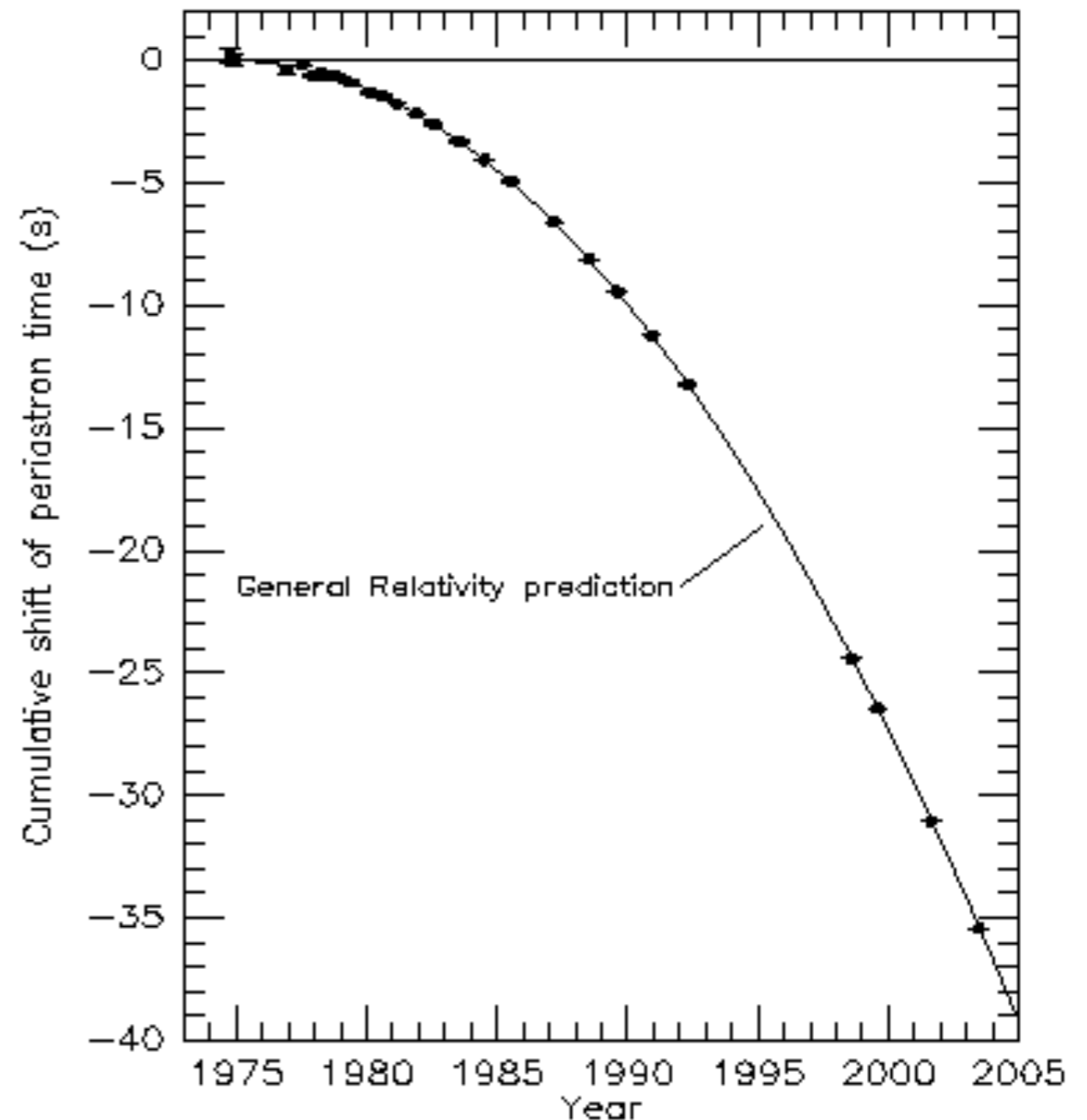
Evidence for Gravitational Waves

- PSR 1913+16 found by Hulse and Taylor in 1974
- It is part of a binary system and the companion is also a neutron star
- Measuring pulse arrival time gives precise measurement of orbital period

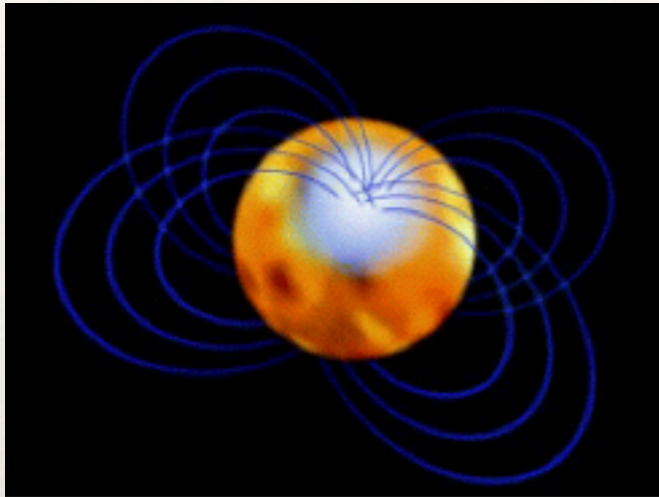


Evidence for Gravitational Waves

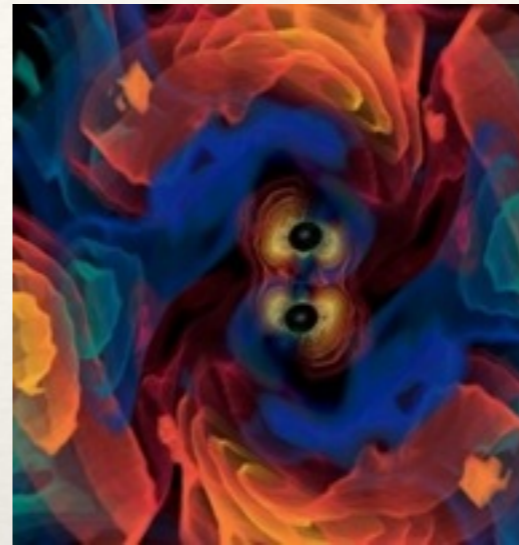
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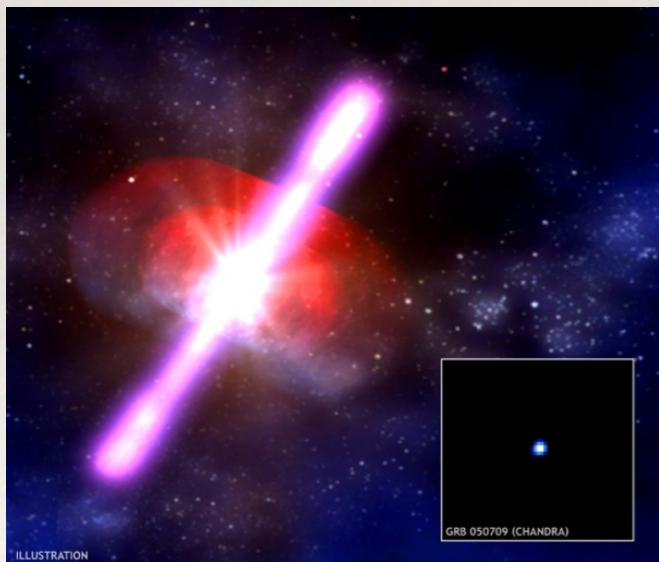
Sources of Gravitational Waves



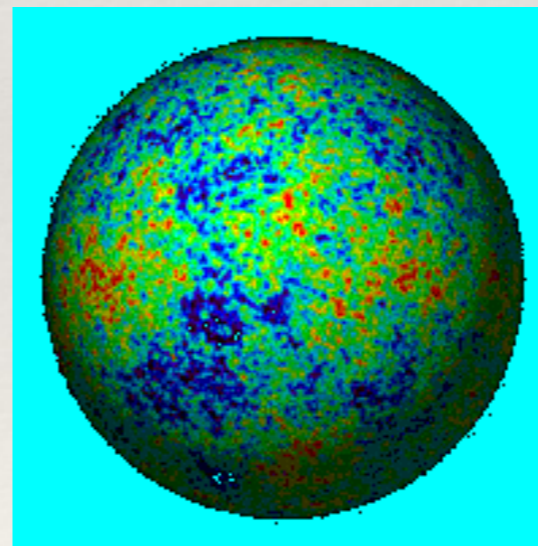
Continuous Waves:
Spinning Neutron Stars



Coalescing Binaries:
Merging black holes and neutron stars

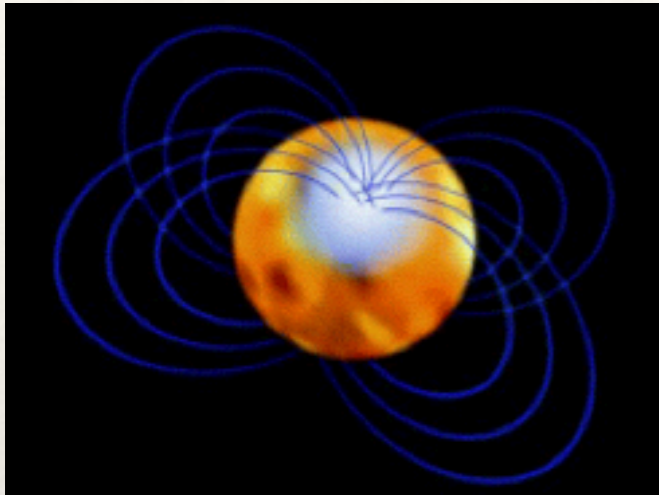


Short Bursts:
Supernovae,
Gamma Ray Bursts



Stochastic Sources:
GW from the Big Bang

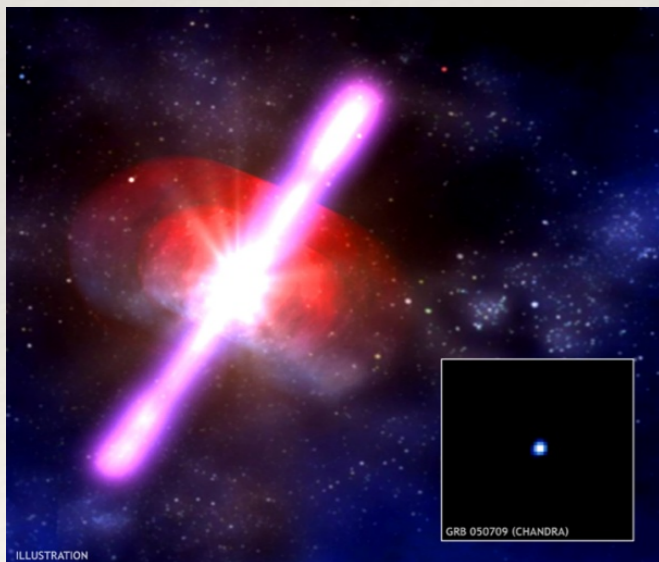
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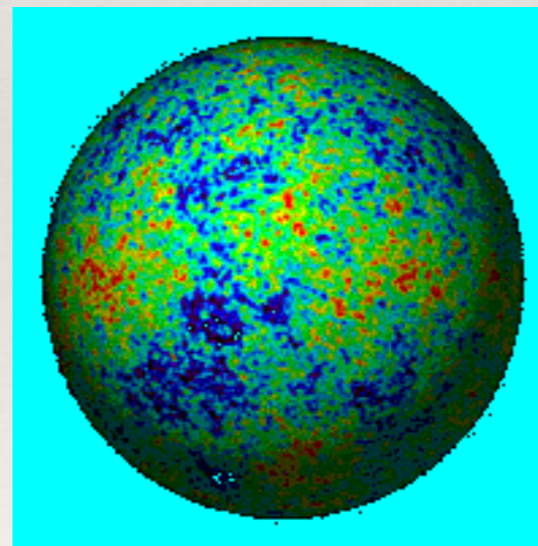
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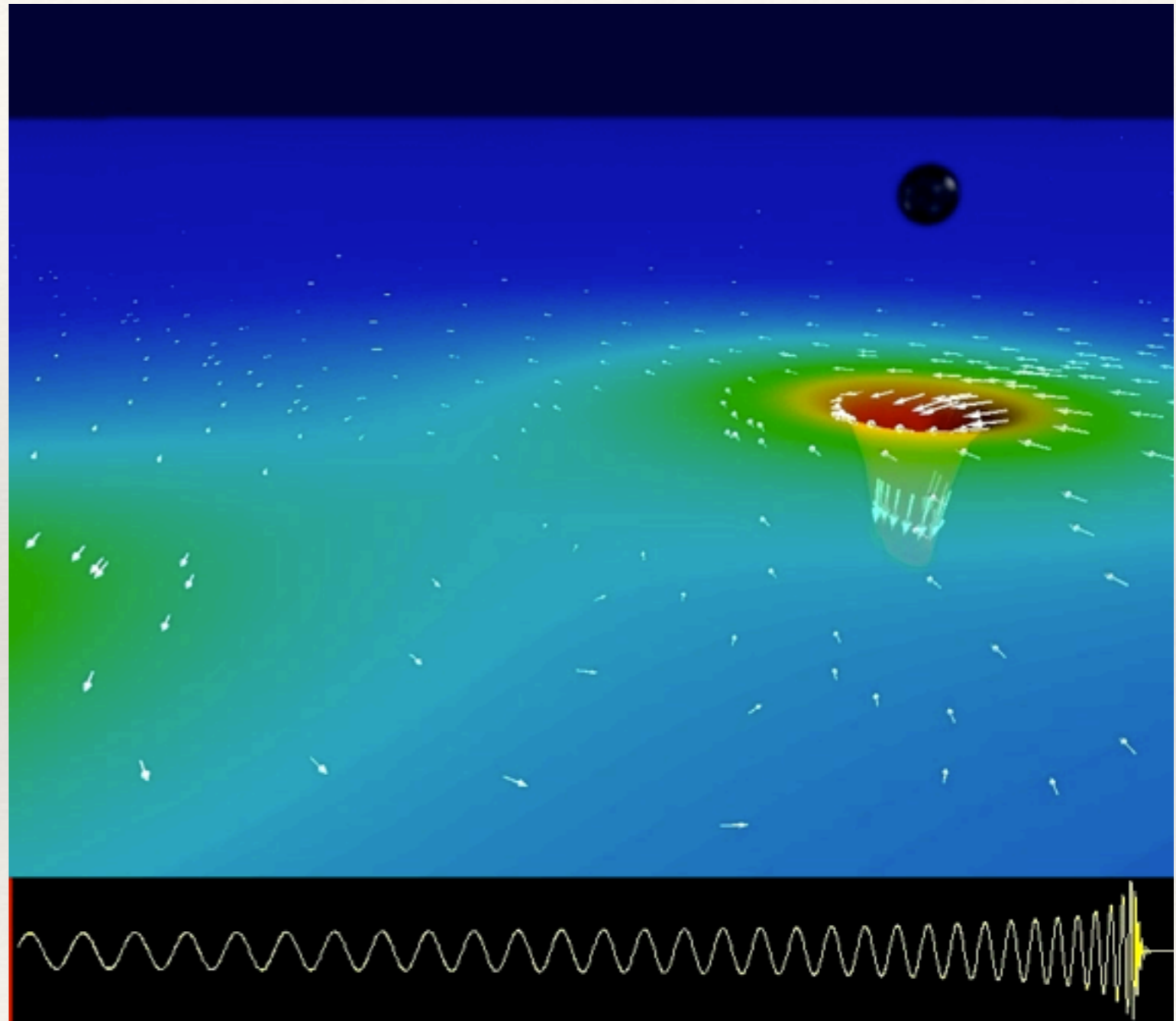
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Stochastic Sources:
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Modelling black hole collisions

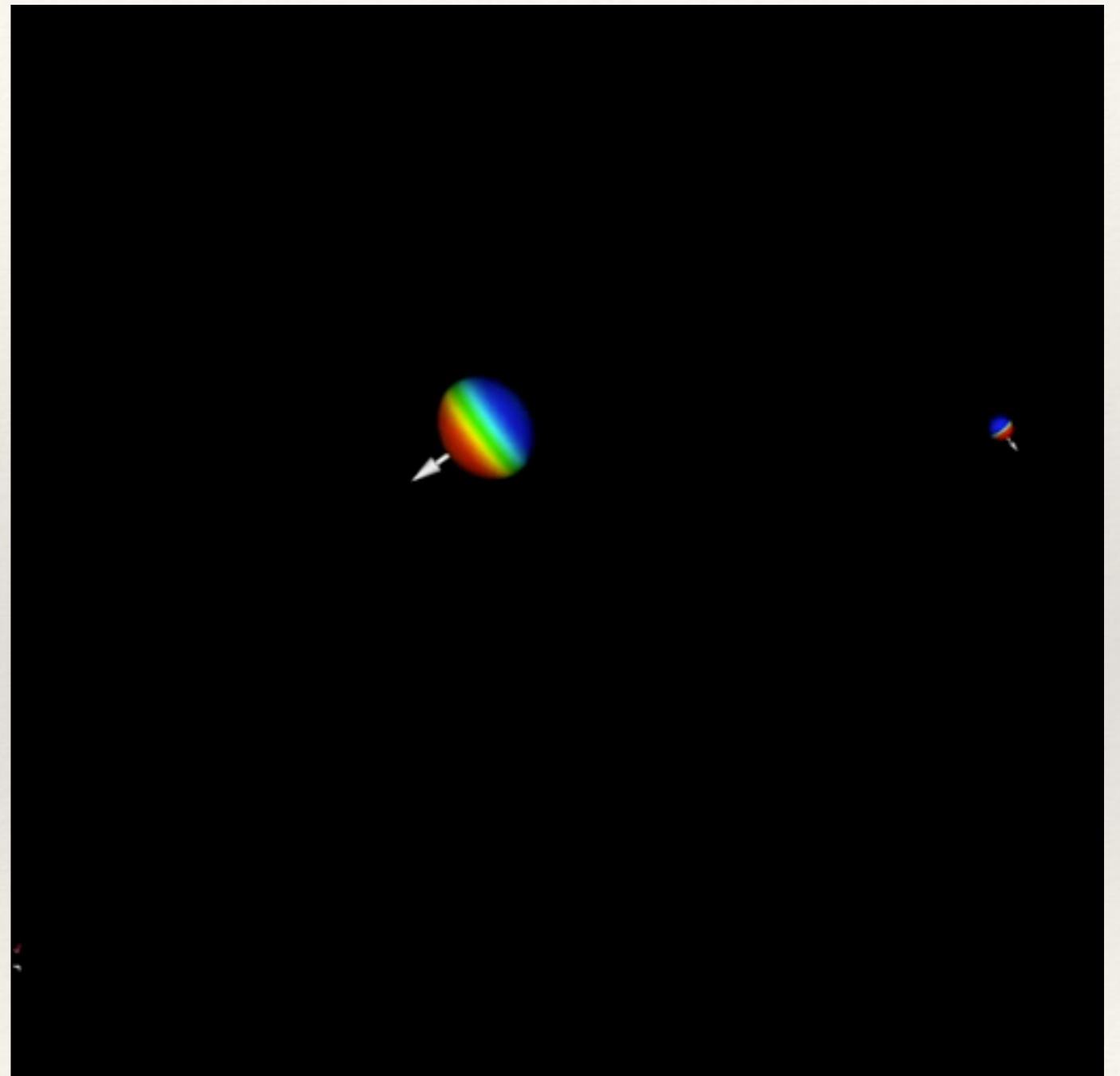
- When black holes are far apart, calculate waveforms analytically
- Close to merger, need to solve non-linear Einstein equations numerically
- Calculate binary evolution and gravitational wave emission



Simulations of Extreme Spacetimes (black-holes.org)

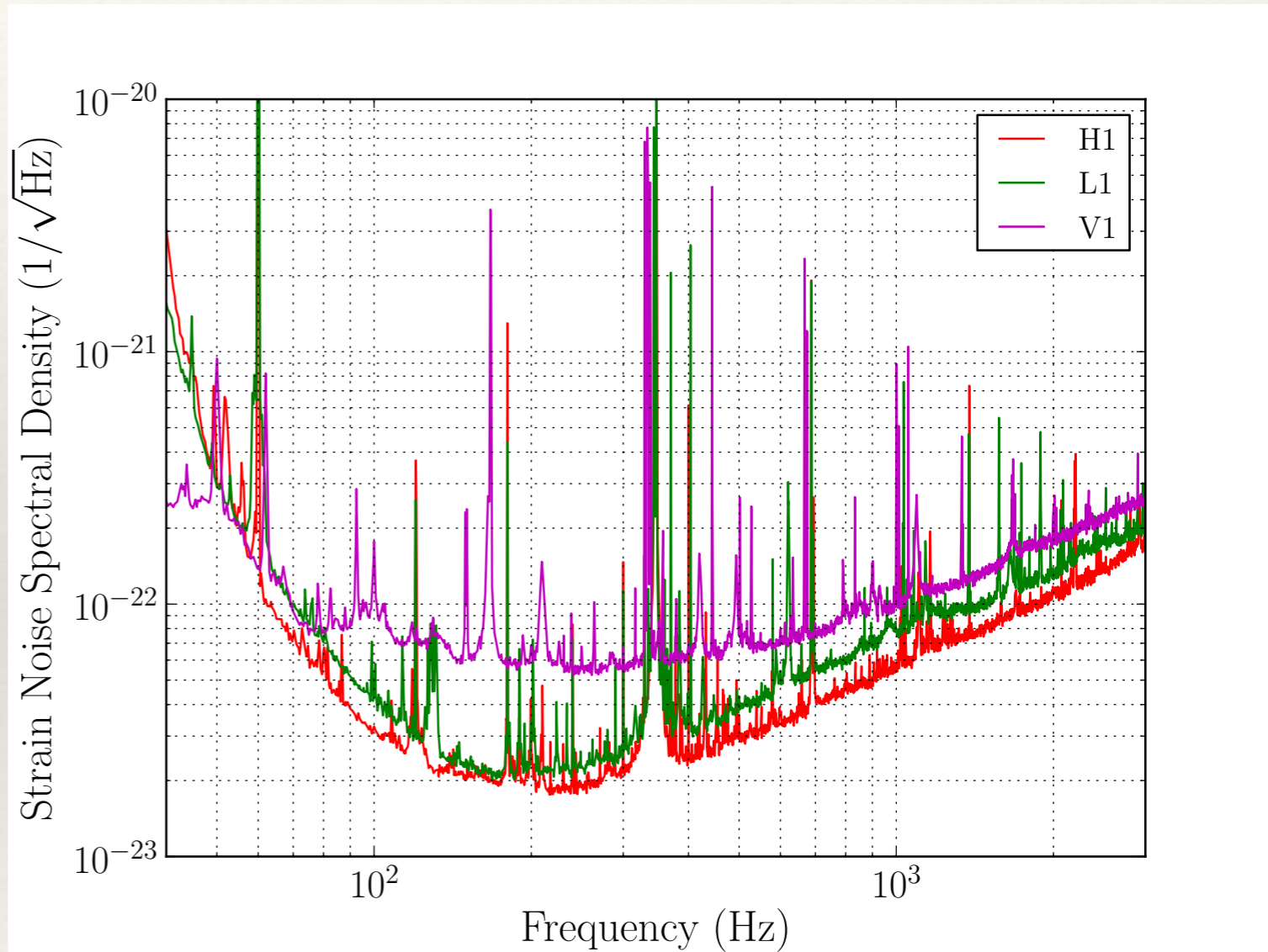
Future Challenges

- Black hole spins complicate matters: cause system to precess
- Naive coverage of parameter space would require $\sim 10,000$ simulations
- Typical simulation takes a month on 256 cores



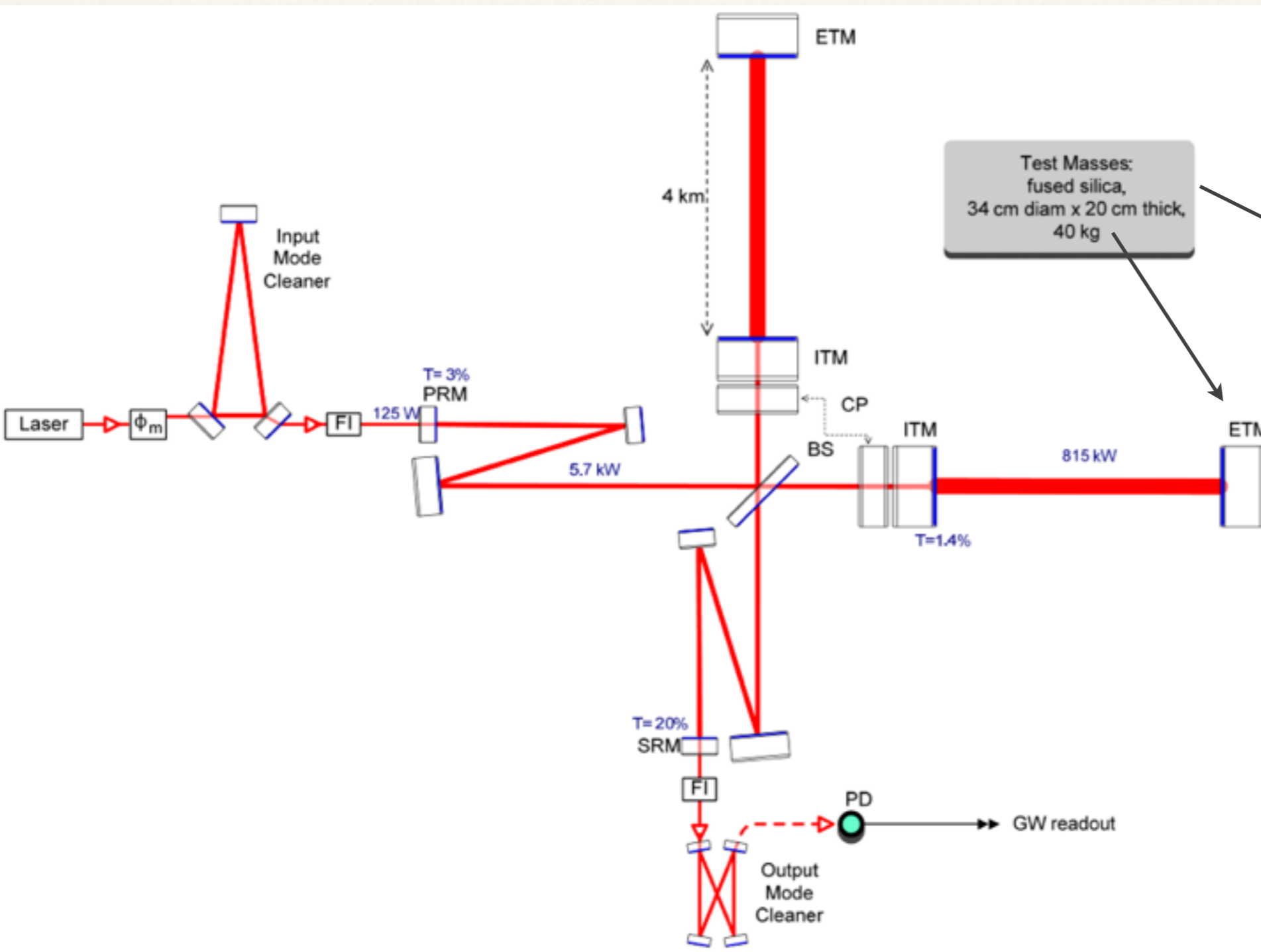
Simulations of Extreme Spacetimes (black-holes.org)

The detectors

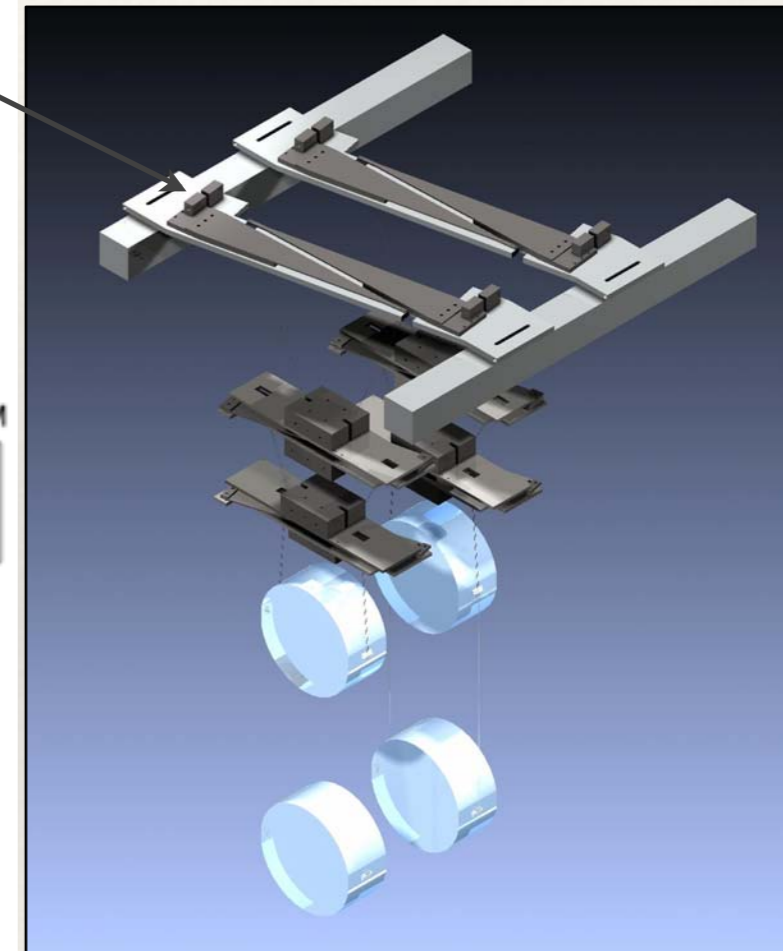


From Abadie et al,
PRD (2012)

Advanced LIGO Detector



Test Masses:
fused silica,
34 cm diam x 20 cm thick,
40 kg



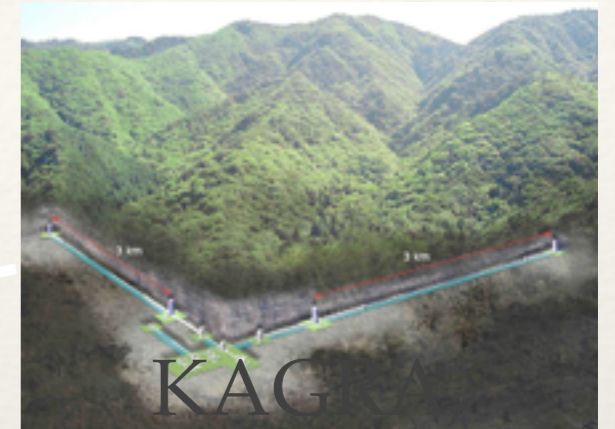
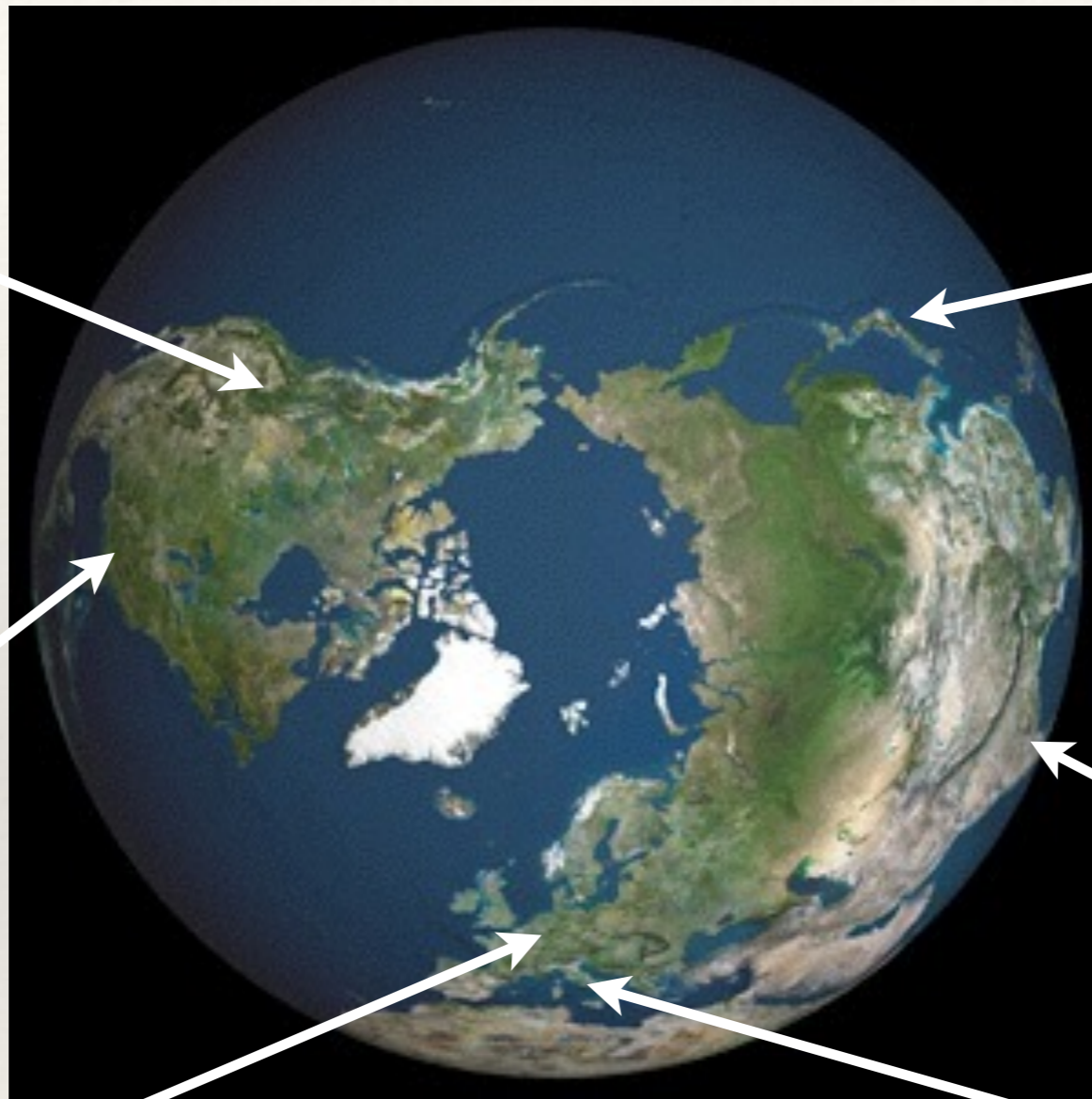
A global network



LIGO Hanford,
USA



LIGO
Livingston,
USA



Japan



LIGO
India



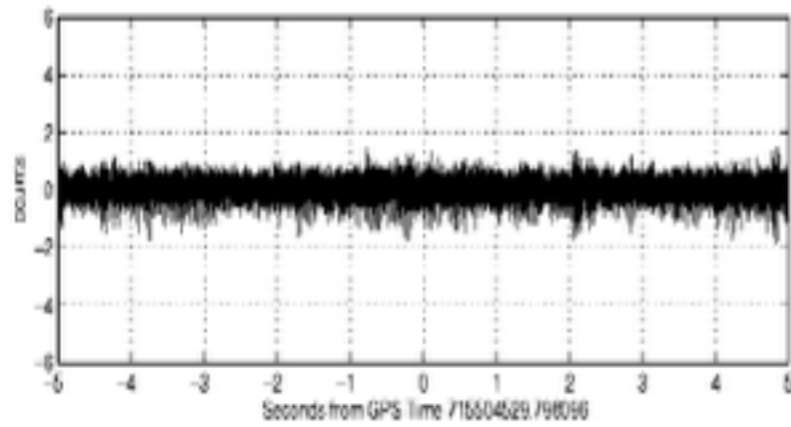
GEO 600,
Germany



Virgo,
Italy

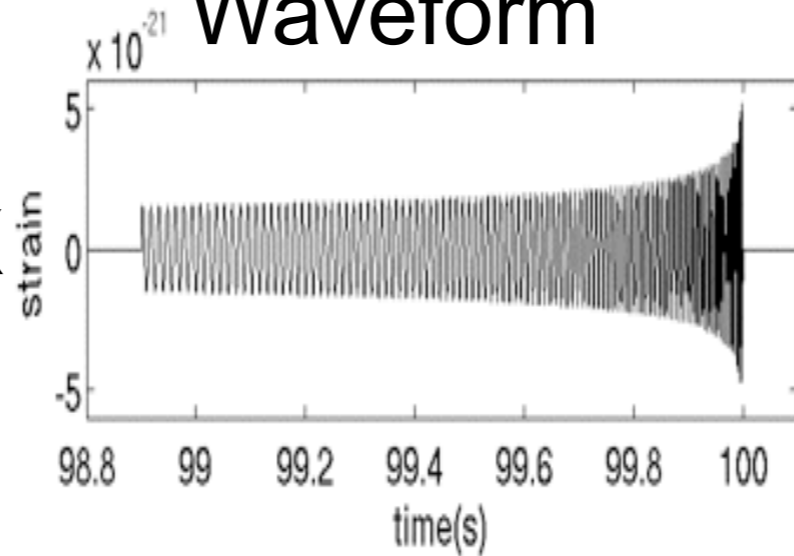
Searching for gravitational waves

Data



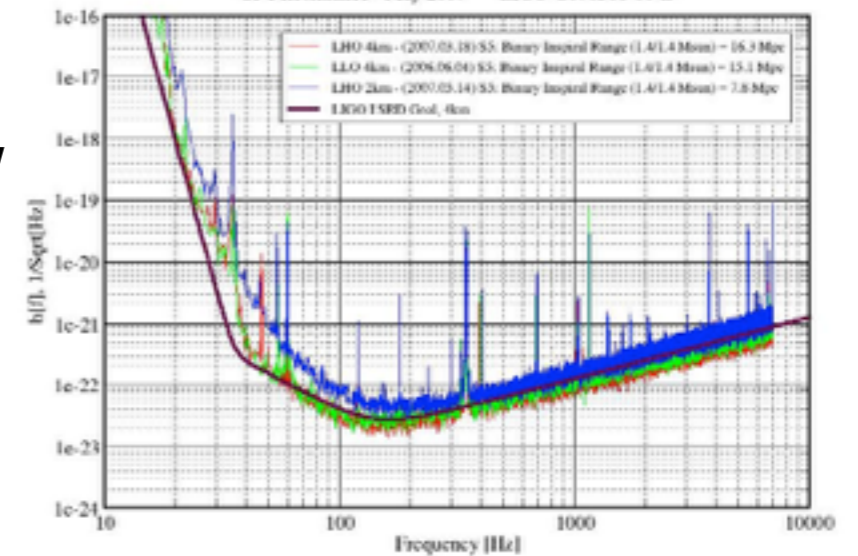
X

Waveform



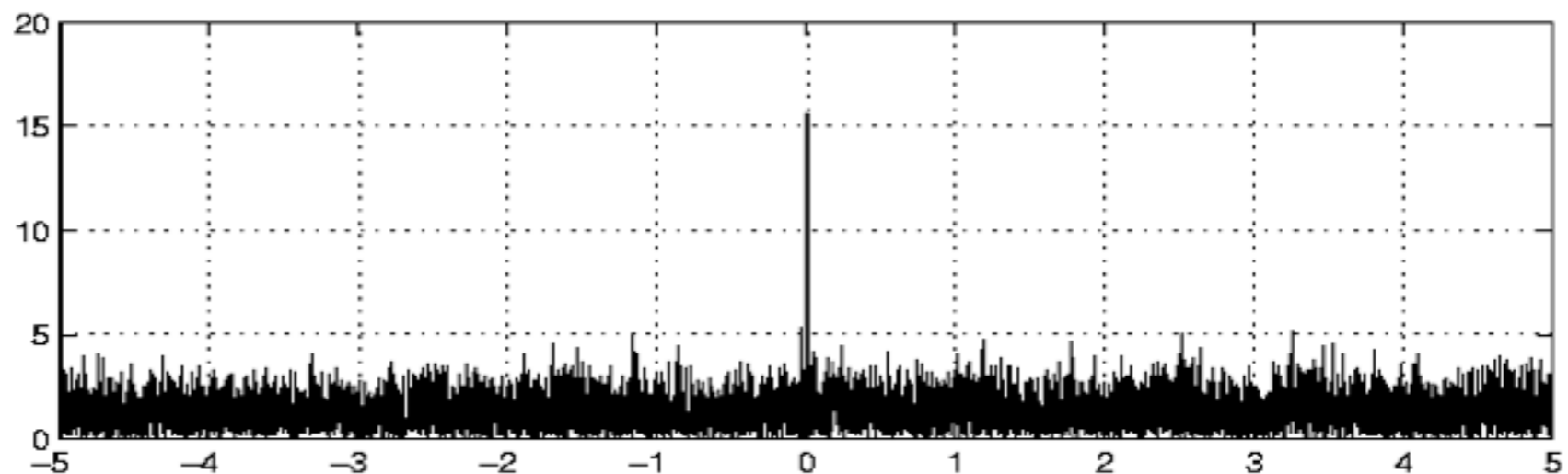
/

Sensitivity



||

SNR



Time of merger

Search results

news

Advanced LIGO

science

students/teachers/public

multimedia

partners

latest news

upcoming events

press releases

press information

past events

blog

"BLIND INJECTION" STRESS-TESTS LIGO AND VIRGO'S SEARCH FOR GRAVITATIONAL WAVES

The LIGO Scientific Collaboration and the Virgo Collaboration completed an end-to-end system test of their detection capabilities at their recent joint collaboration meeting in Arcadia, CA. Analysis of data from LIGO and Virgo's most recent observation run revealed evidence of the elusive signal from a neutron star spiraling into a black hole. Shortly after the collaboration approved a scientific paper reporting the ground-breaking "discovery", LIGO and Virgo management revealed that the signal was a "blind injection" --- a fake signal secretly added to the data to test the detector and analysis.

While the scientists were disappointed that the discovery was not real (they knew that it could be a blind injection), the success of the analysis was a compelling demonstration of the collaboration's readiness to detect gravitational waves. LIGO and Virgo scientists are looking forward to observations with the advanced detectors which are expected to contain many real signals from the distant reaches of the universe.

GRAVITATIONAL WAVES:

There's a lot at stake here. [Gravitational waves](#), a firm prediction of Einstein's General Theory of Relativity, have never been directly detected, although there is convincing indirect evidence for their existence from [precise timing of the orbits of binary pulsars in the Galaxy](#). The direct detection of these waves through the tiny distortions of space-time that they produce when they arrive at Earth from distant astrophysical sources would be a major scientific milestone, and would open up the new field of gravitational-wave astronomy.

CATCH THE

CATCHING

GRAVITATIONAL

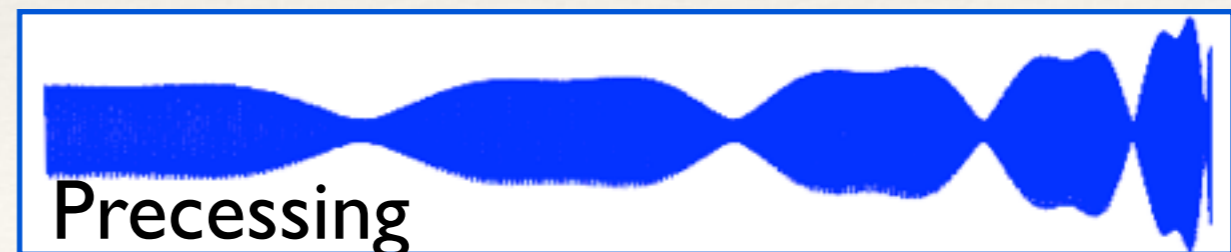
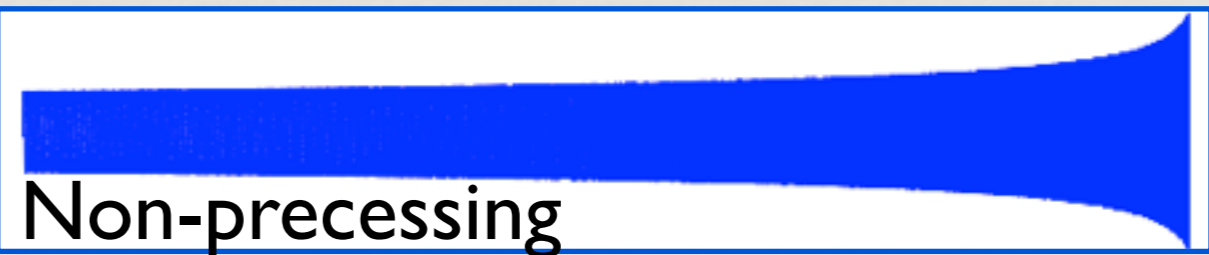
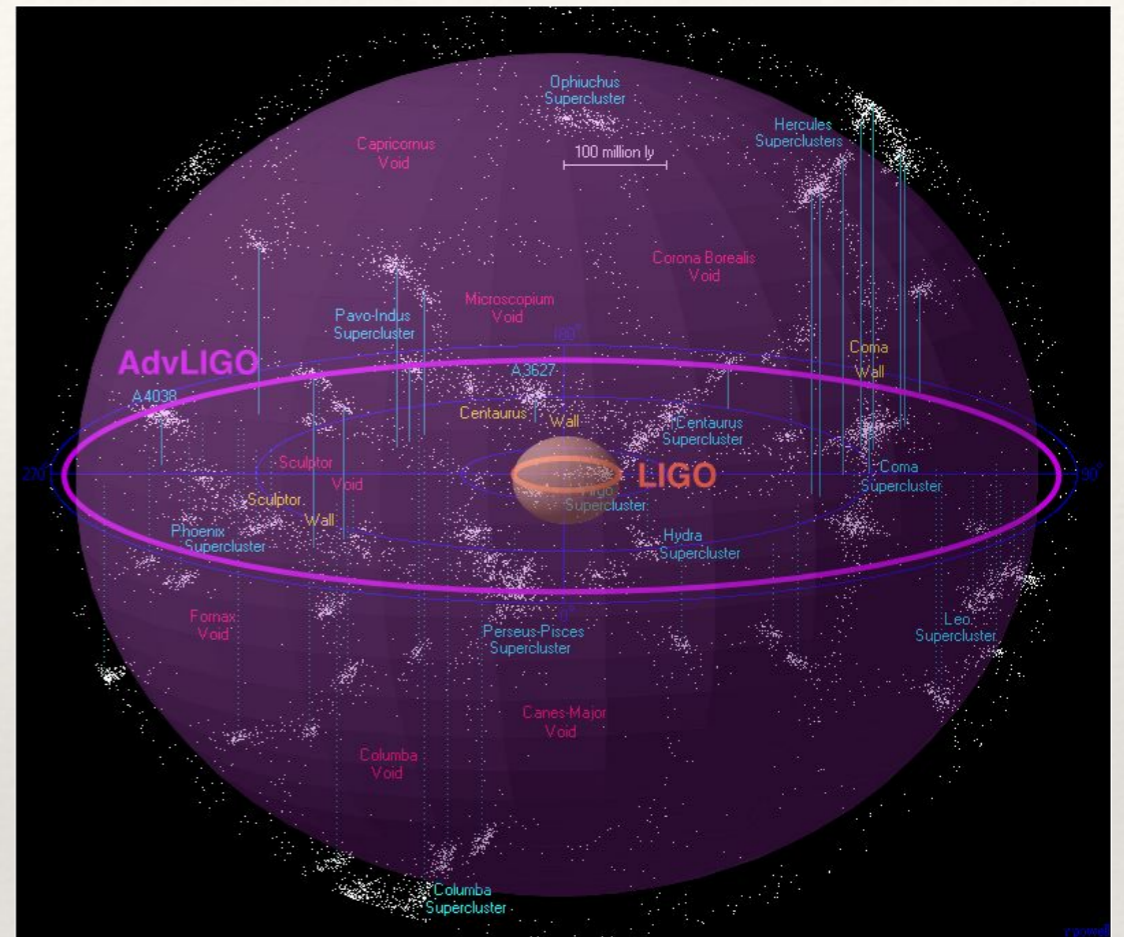
AS A PROBE O

LIGO

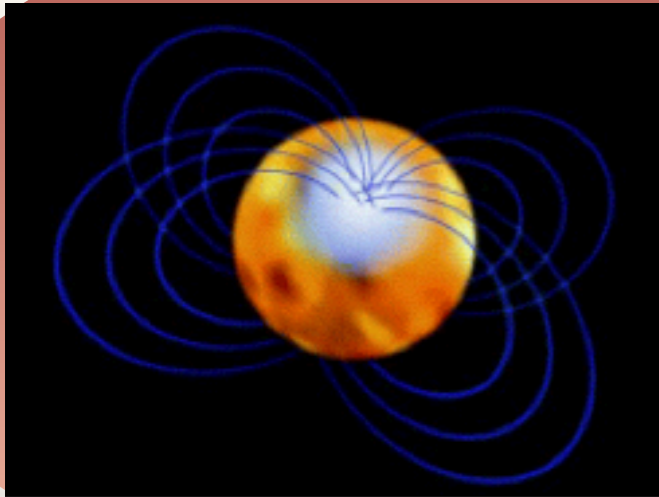


Future challenges

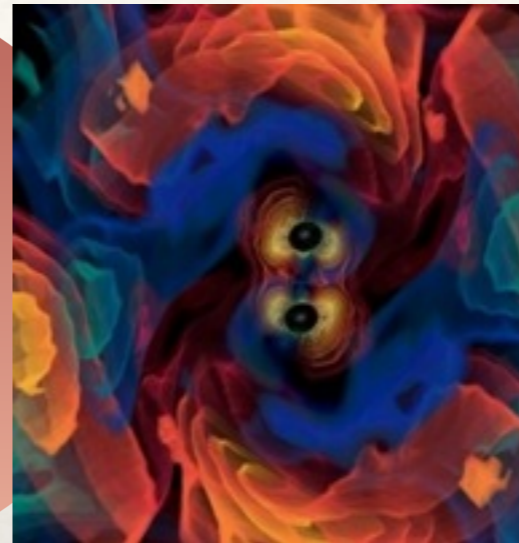
- Advanced detectors will have ten fold increase in sensitivity, expect to make first observations
- Need to search months or years of data using $\sim 10,000$ template waveforms for early runs (2015-16)
- Increases to millions of waveforms at final sensitivity (~ 2019), incorporating spin and precession effects
- Data is full of noise transients (“glitches”) that can mask signals



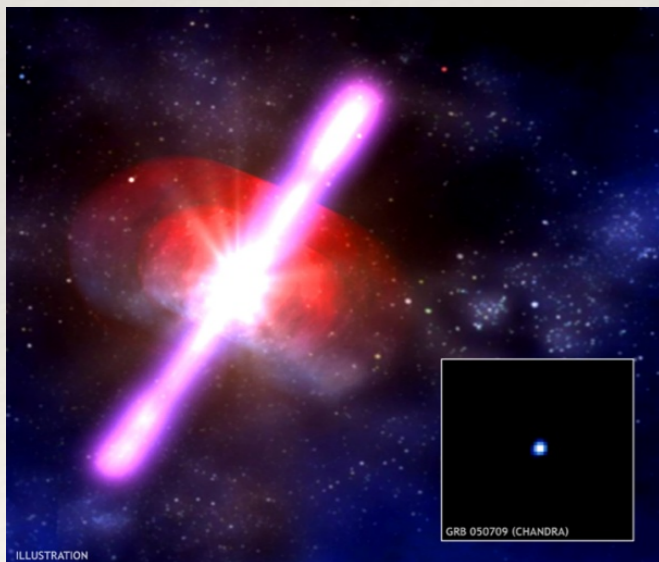
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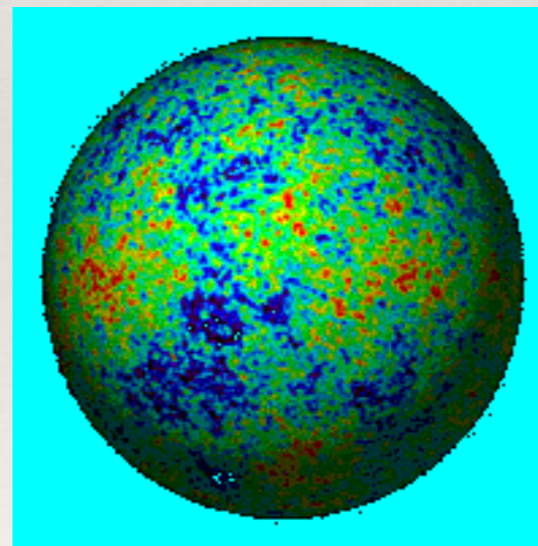
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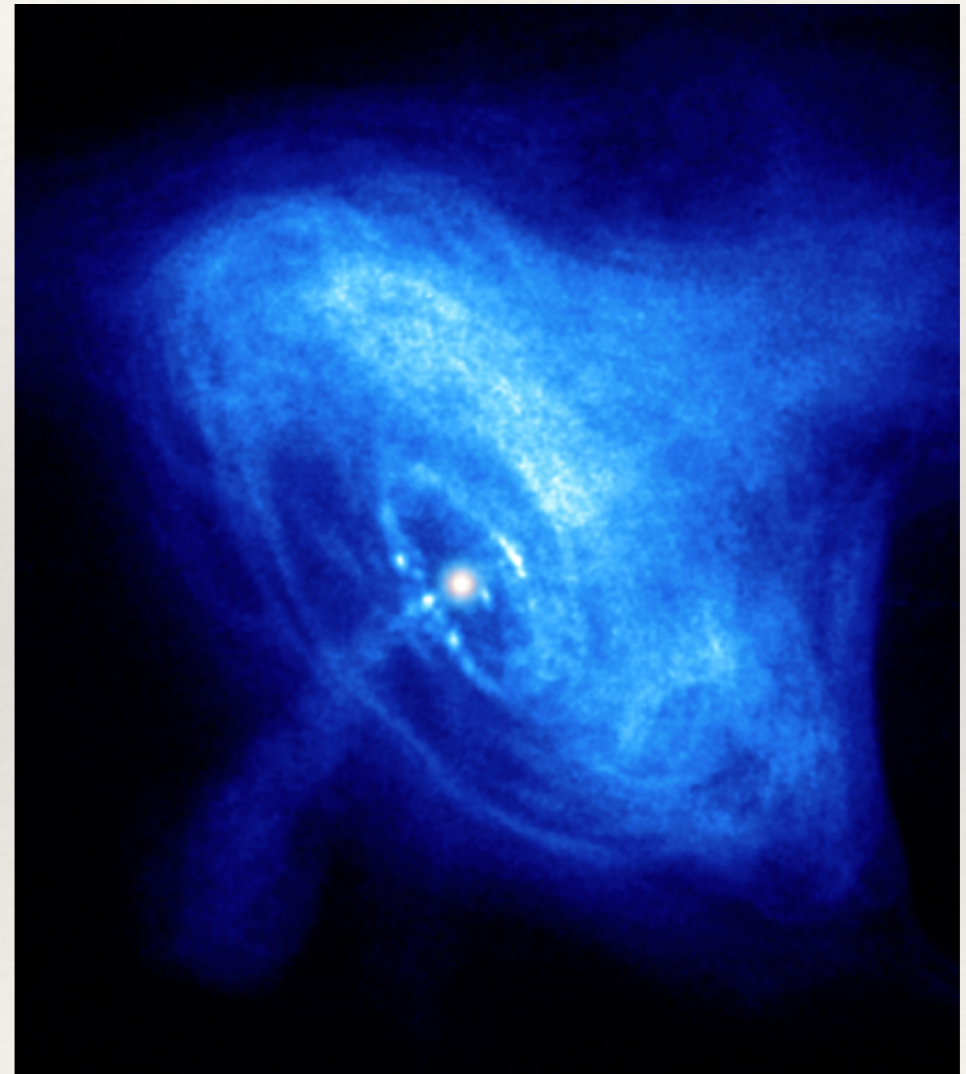
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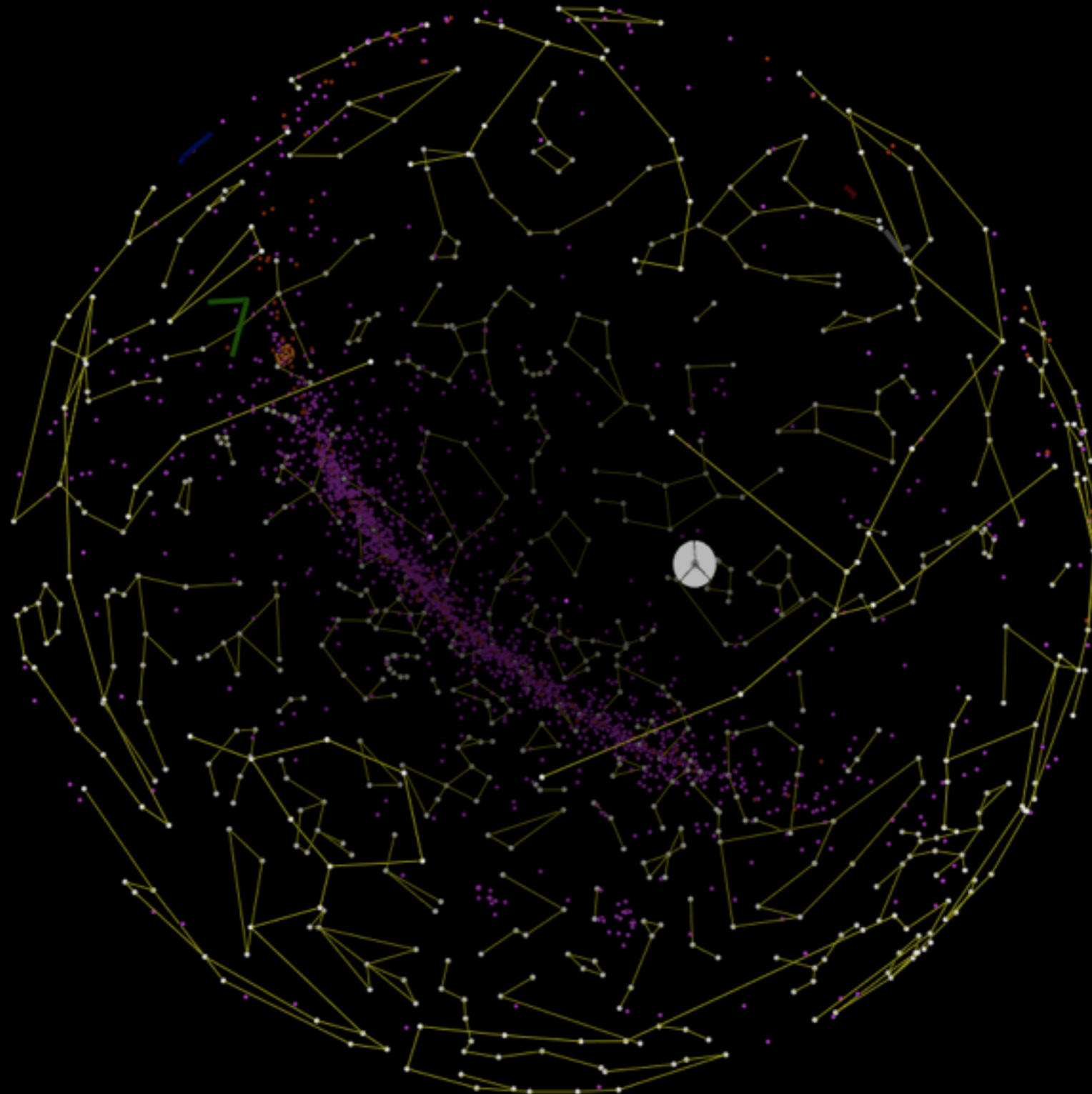
- Small “mountain” on a spinning neutron star gives a continuous gravitational wave signal
- Motion of earth modulates signal frequency
- Fully coherent search requires more computing than available



Einstein@Home

Einstein@Home

International Year of Astronomy 2009



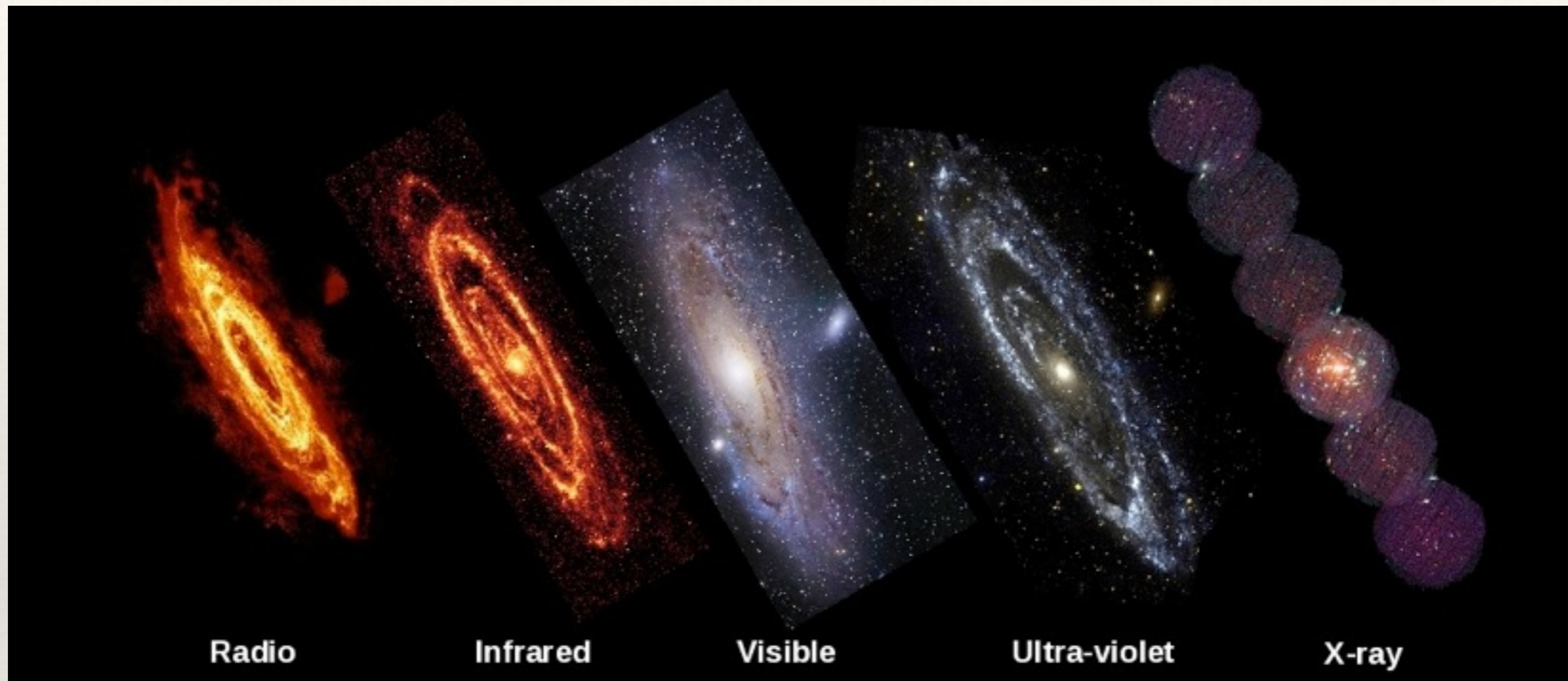
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Team: Albert-Einstein-Institut Hannover (AEI)
Project Credit: 2342206.43
Project RAC: 1429.52
WU Completed: 24.40 %
WU CPU Time: 00:51:00

Search Information

Ascension: 299.14 deg
Declination: 29.79 deg
DM: 134.40 pc/cm³
Orb. Radius: 0.155 ls
Orb. Period: 697 s
Orb. Phase: 4.36 rad

The payoff



- New insights provided by “multi-wavelength” astronomy
- Gravitational wave observations will provide “multi-messenger” astronomy